

The Cosmic-Ray Air Shower Signal in Askaryan Radio detectors

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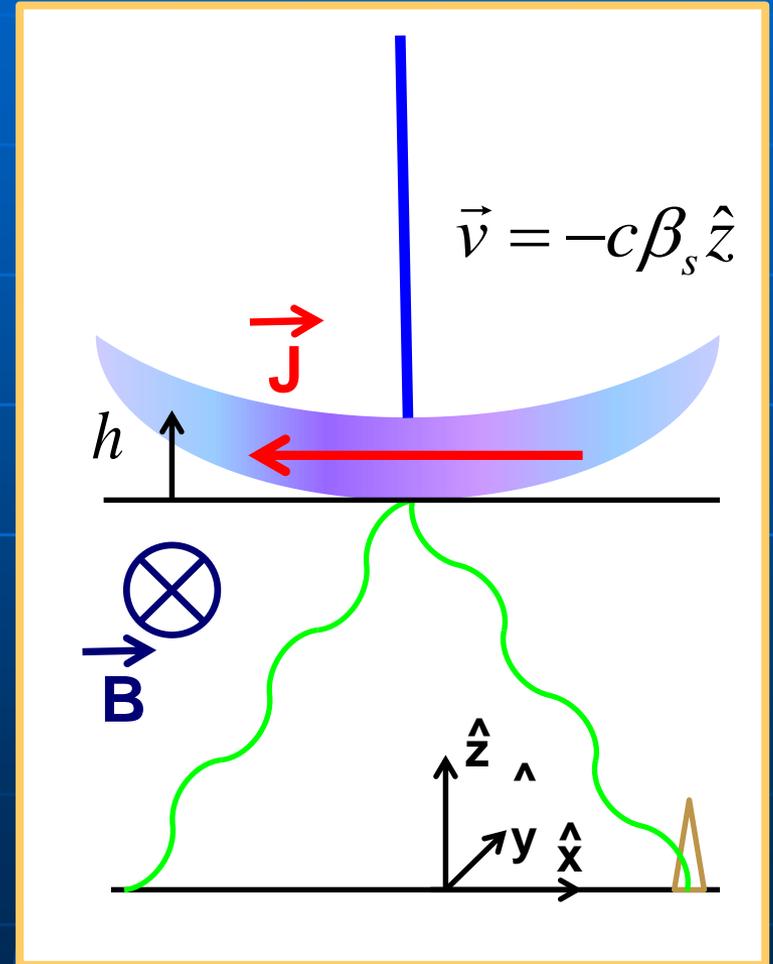
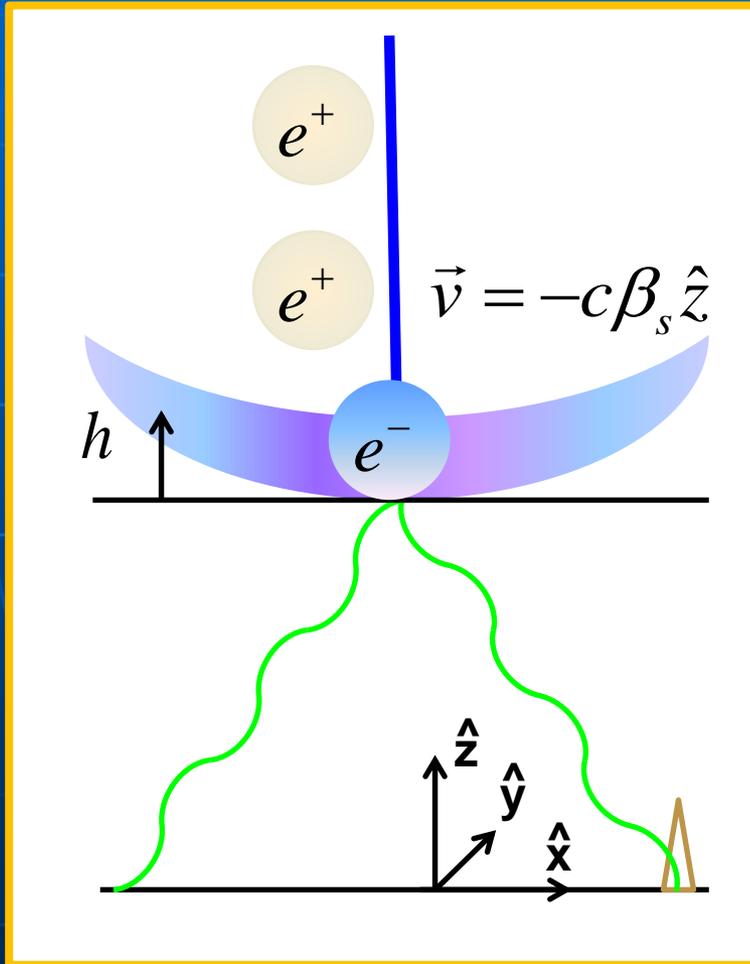
IIHE

VUB, ULB, KVI

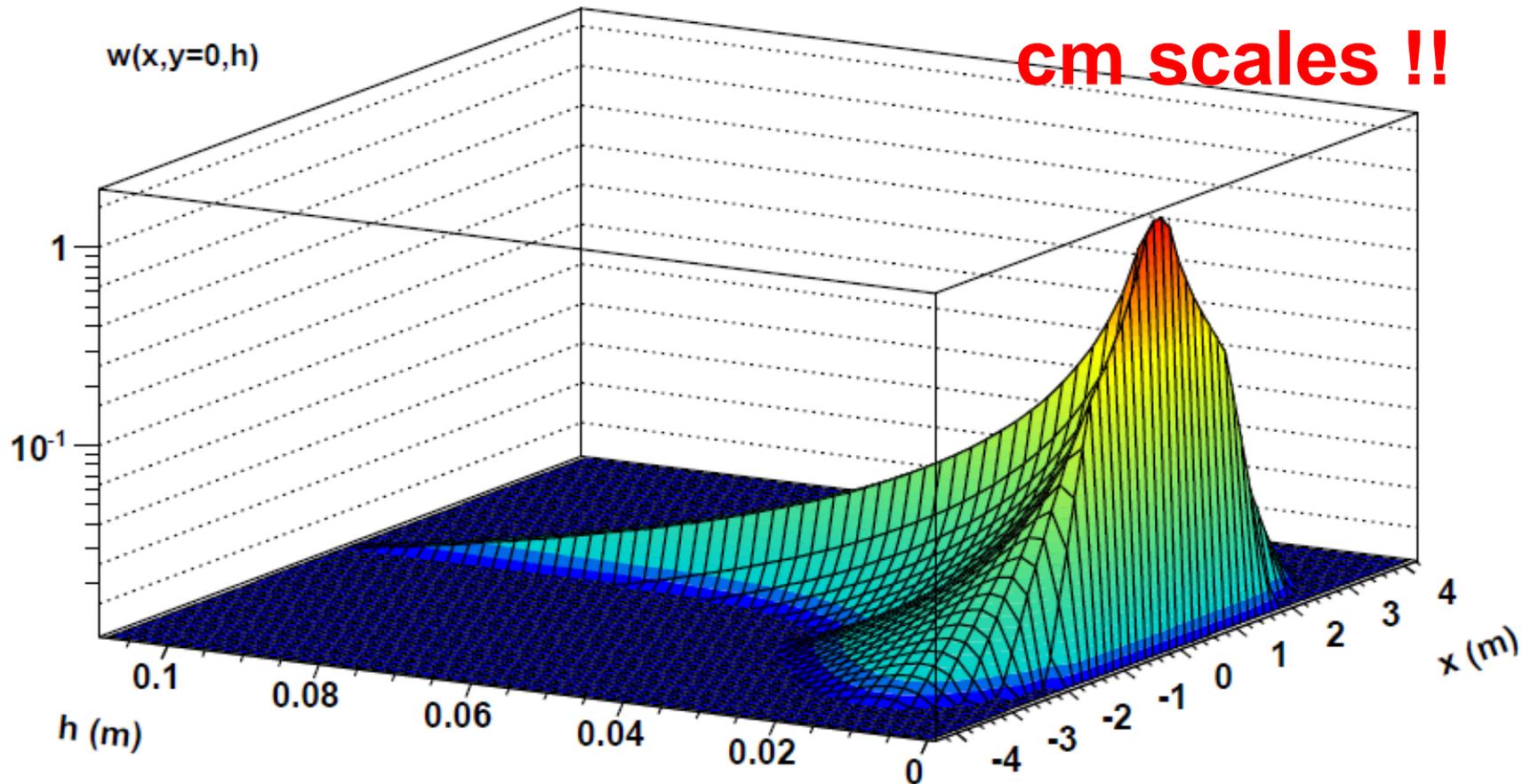


Radio Emission Mechanisms:

Charge excess: Geomagnetic:



CX-MC-GEO+MCFIT: The particle distributions in the shower front



Emission Mechanisms

From Currents to radiation.

$$A^\mu(\vec{x}, t) = \frac{J^\mu(t')}{|D(\vec{x}, t)|}$$
$$D = R(1 - n\beta \cos(\theta))$$
$$= R \frac{dt}{dt'}$$

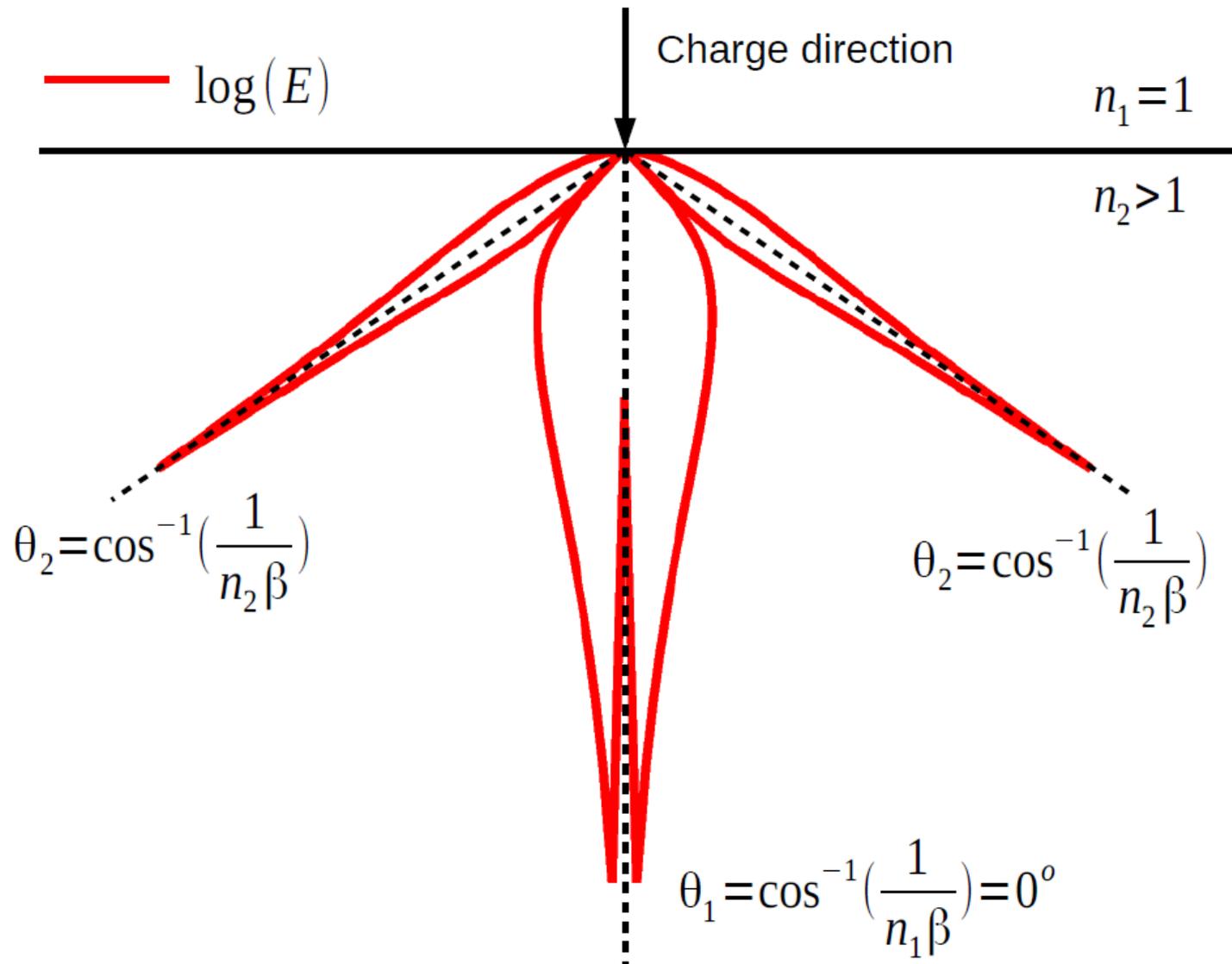
$$\vec{E}(\vec{x}, t) =$$
$$-\frac{d}{dt} \vec{A}(\vec{x}, t) - \frac{d}{d\vec{x}} A^0(\vec{x}, t)$$

$$\vec{E}(\vec{x}, t) \propto \frac{1}{D^2}$$

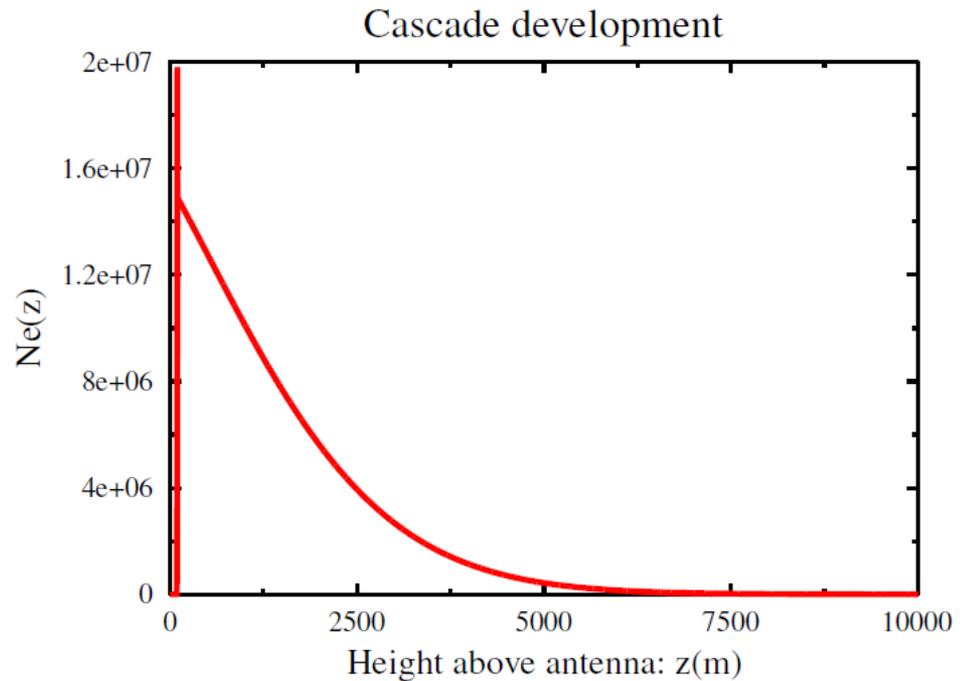
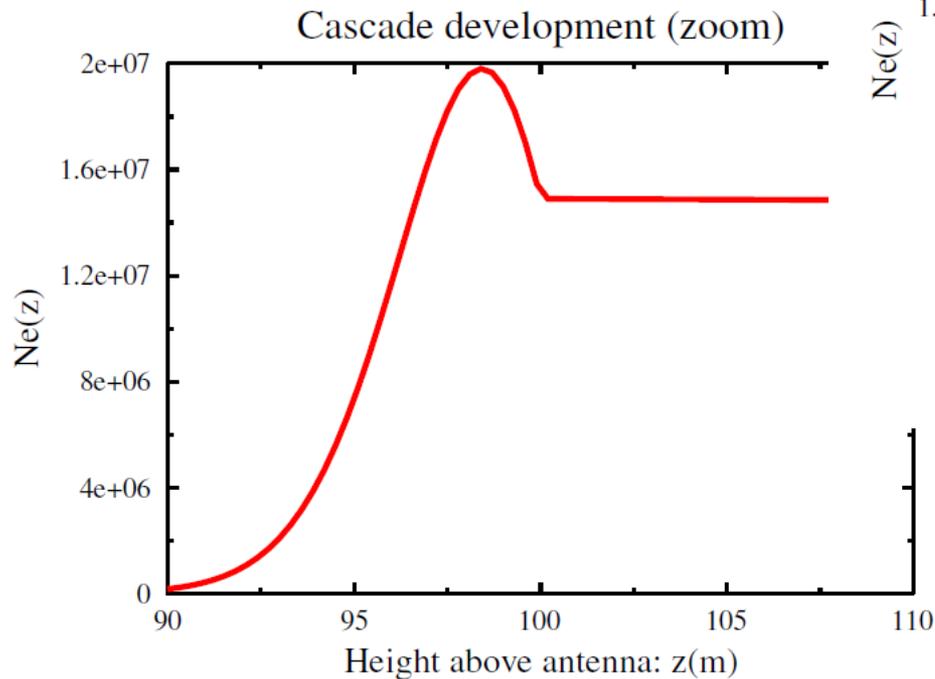
D can vanish for realistic cases,
 $n = n(z) \neq 1 \rightarrow$ Cherenkov !

[ArXiv: 1503.02808](https://arxiv.org/abs/1503.02808)

Recent work: Coherent Transition Radiation

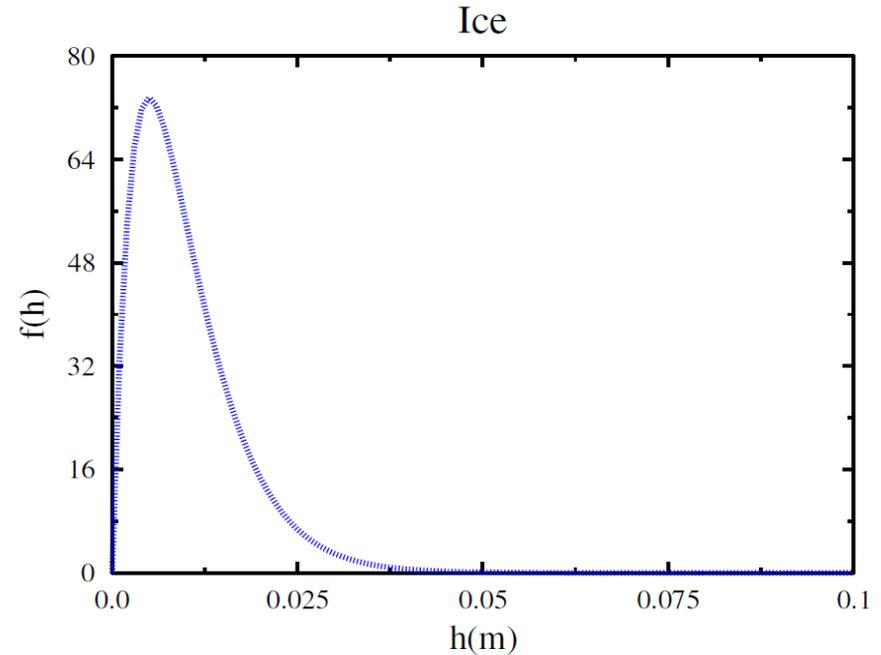
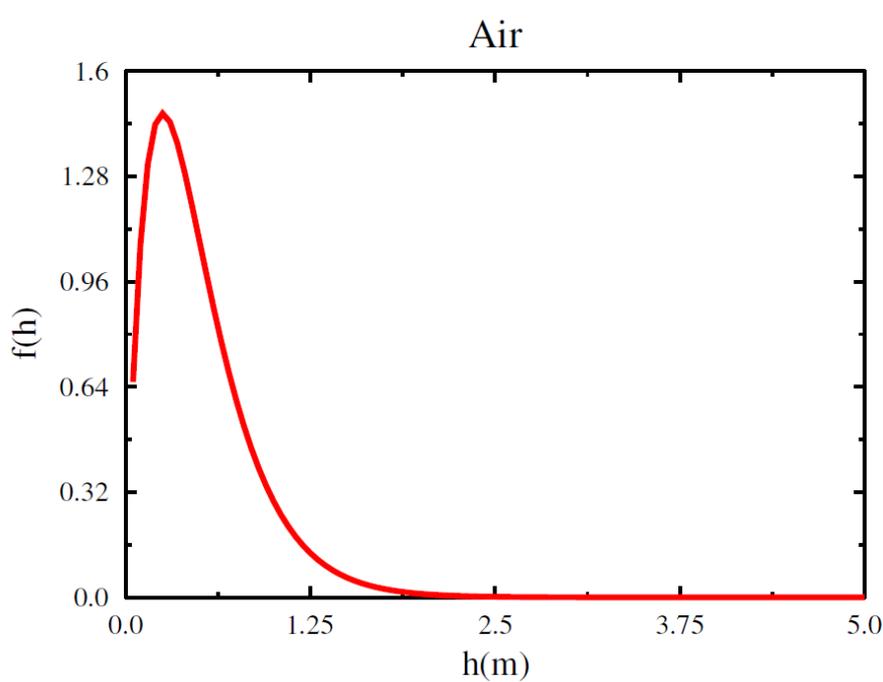


The cosmic-ray air shower hitting an ice surface



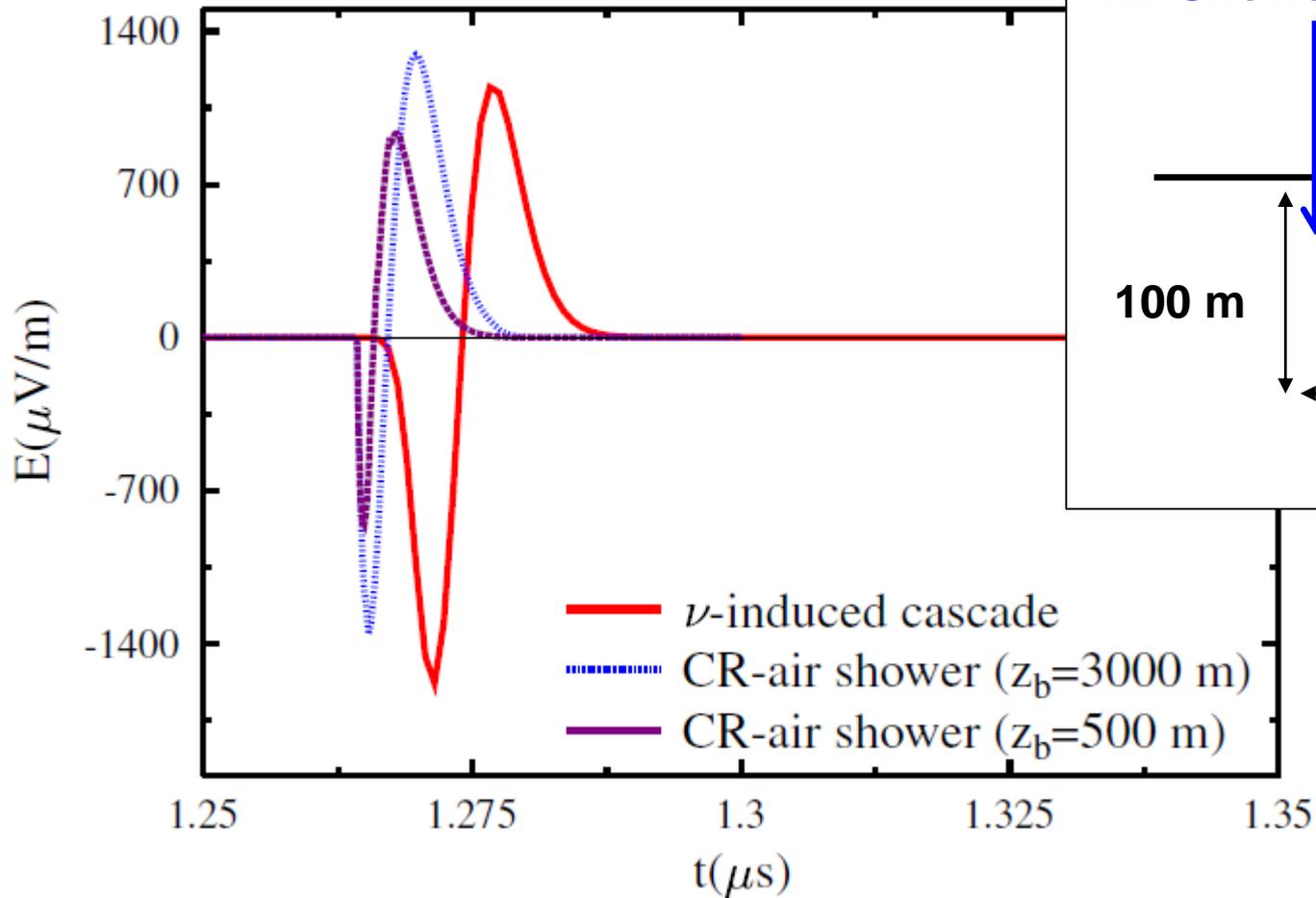
The cosmic-ray air shower hitting an ice surface

Effective parameterization of the air shower front.



Lateral dimension folded onto longitudinal dimension. **Full 3D description is under development.**

The air shower signal vs the neutrino induced cascade



Air Shower

Neutrino induced cascade

100 m

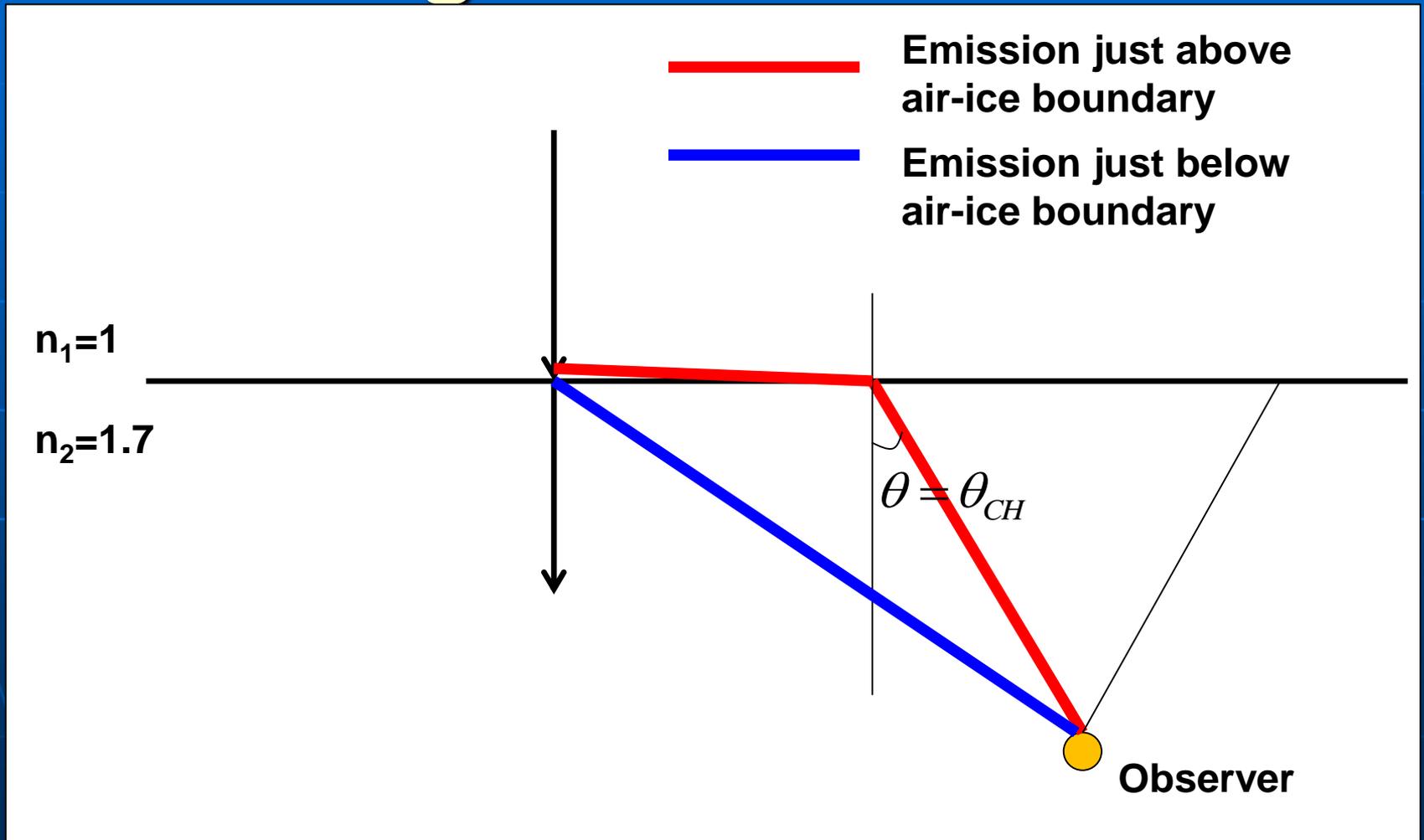
250 m

Observer

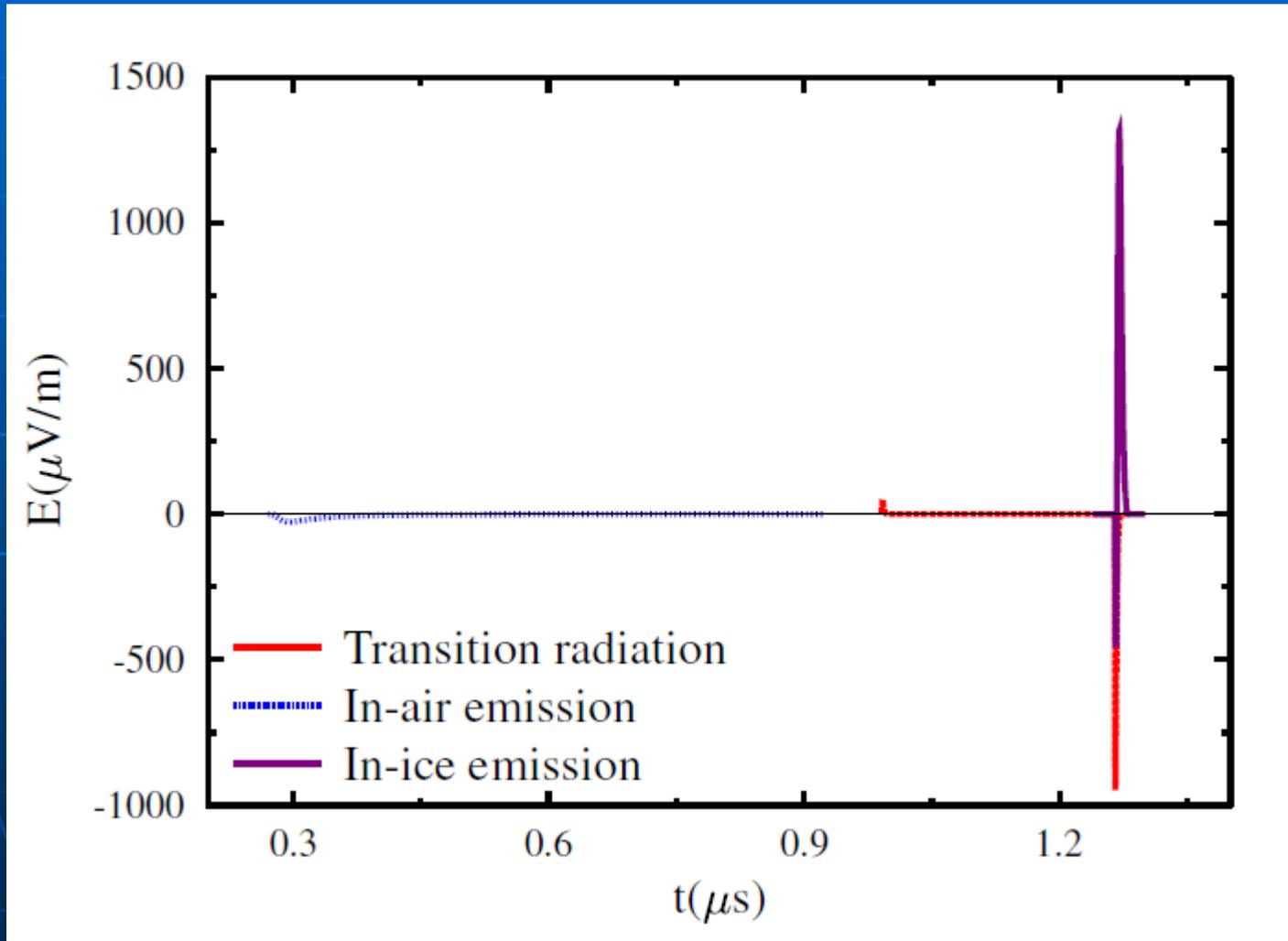
What can we learn from the Cosmic-Ray air shower signal?

- The in-ice emission is due to the same mechanism (Askaryan emission) as for a neutrino induced cascade. **The expected signal is similar for CR and neutrino induced cascades.**
- In combination with surface detector the detected signal might provide an **excellent (cross-) calibration.**
- The in-ice Askaryan radio detector can be used as a **cosmic-ray air shower** detector.

The cosmic-ray air shower hitting an ice surface



The cosmic-ray air shower hitting an ice surface



Conclusions and Outlook

- The cosmic-ray air shower signal in Askaryan radio detectors has been modeled giving:
 - 1) The in-air emission solving for refraction.
 - 2) The transition radiation at the air-ice surface.
 - 3) The in-ice emission.
- The signal is comparable to that of a neutrino induced cascade and should thus be measurable. Possible applications are:
 - 1) (Cross-) calibration in combination with a surface detector.
 - 2) Air shower physics
 - 3) On-site feasibility of the detection technique.
- A more detailed modeling of the shower front is **under development.**

Recent work: Coherent Transition Radiation

$$\vec{E}(\vec{x}, t) = \frac{ex^i}{4\pi\epsilon_0 c} \lim_{\epsilon \rightarrow 0} \left(\frac{1}{D^2_{z_B + \epsilon}} - \frac{1}{D^2_{z_B - \epsilon}} \right) \delta(z - z_B)$$

The delta-function can be resolved by considering coherent emission over a macroscopic particle distribution!!!

$$\vec{E}(\vec{x}, t) = \frac{eN_e(t_r) f(h) x^i}{4\pi\epsilon_0 c} \left(\frac{1}{D^2_{z_B + \epsilon}} - \frac{1}{D^2_{z_B - \epsilon}} \right)_{h=ct_r - z_B}$$

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