An estimate of the spectral intensity expected from molecular Bremsstrahlung radiation in extensive air showers

Imen Al Samarai

Institut de Physique Nucléaire d'Orsay - CNRS/IN2P3

11 juin 2014

◆□▶ ◆□▶ ◆三▶ ◆三▶ 三三 のへで

Outline

Introduction

2 Ionization electrons along the shower track A crude model for EAS Production of ionization e⁻ Time evolution of ionization e⁻ flux

(日) (四) (코) (코) (코)

- 4 Simulation results
- 5 Conclusion & Outlook

Introduction

Ionization electrons along the shower track $\mu-{\rm wave}$ emission from MBR Simulation results Conclusion & Outbook

Motivation

Depict analytically the MBR mechanism

- Evaluation of the number of primary charged particles in an EAS
- Production of ionization electrons and their time evolution
- MBR emission using the free-free approach



A crude model for EAS Production of ionization e^- Time evolution of ionization e^- flux

Total number of primary e^+/e^- per unit surface

$$n_{e,p}(r,a) = N_{e,p}(a) \frac{ldf(r,a)}{2\pi \int dr \ r \ ldf(r,a)}$$

- Total number of primaries at altitude *a* using Gaisser-Hillas formula
- NKG function describing the lateral distribution (LDF)



In this study, parameters are tuned to apply to a vertical shower with energy $10^{17.5}$ eV to compare with [1].

A crude model for EAS **Production of ionization e** Time evolution of ionization e⁻ flux

• □ ▶ • • □ ▶ • • □ ▶ •

« plasma »

hower front

- Presence of a weakly ionized plasma
- Number of ionized electrons per unit length per energy band

$$\frac{d^2 N_{e,i}}{da \ dT_e}(a, T_e) = \rho_m(a) \quad f_0(T_e) \quad \left\langle \frac{dE}{dX} \right\rangle \quad \frac{1}{I_0 + T_e}.$$

with parametrization from [2] :

$$f_0(T_e) = \frac{K}{1 + (T_e/\overline{T})^{2.1}}$$

is the distribution in kinetic energy of the resulting ionization electrons and $\left\langle \frac{dE}{dX} \right\rangle$ the mean energy loss of primary electrons per grammage unit

A crude model for EAS **Production of ionization e** Time evolution of ionization e⁻ flux

(日) (同) (三) (



Instantaneous flux of ionization electrons per kinetic energy band

$$\phi_{e,i}^0(r,a,T_e) = \frac{c\beta(T_e)f_0(T_e)}{2(I_0+T_e)} \left\langle \frac{dE}{dX} \right\rangle \ \rho_m(a) \ n_{e,p}(r,a).$$

< ロ > < 同 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ >

- Time evolution of the flux of ionization electrons fully encompassed in the energy distribution time evolution $f(T_e,t)$
- Boltzmann equation accounting for all interactions at work

Assumptions :

- Ionization electrons static in space (low energy electrons, rate of disappearence ${\sim}100~\rm{ns})$
- Neglect absorption effects



A crude model for EAS Production of ionization e⁻ Time evolution of ionization e⁻ flux

$$\frac{\partial f}{\partial t}(T_e, t) = -n_m(a)c\beta(T_e) \bigg(\sigma_{\rm att}(T_e) + \sigma_{\rm exc}(T_e) + \sigma_{\rm ion}(T_e)\bigg) f(T_e, t)$$
$$+ n_m(a)c \int_{T_e}^{T_e^{\rm max}} dT'_e \beta(T'_e) \bigg(\frac{d\sigma_{\rm ion}}{dT_e}(T'_e, T_e) + \frac{d\sigma_{\rm ion}}{dT_e}(T'_e, T'_e - T_e)\bigg) f(T'_e, t)$$

+ $n_m(a)c \int_{T_e}^{T_e^{max}} dT'_e \beta(T'_e) \frac{d\sigma_{exc}}{dT_e}(T'_e, T_e) f(T'_e, t)$



With σ_i being the cross-sections of interest and $n_m(a)$ the density of molecules at altitude a

A crude model for EAS Production of ionization e⁻ Time evolution of ionization e⁻ flux

- Rates of collision $\tau_i^{-1} = n_m(a)c\beta(T_e)\sigma_i(T_e)$
- Process description of attachment, ionization and excitation from [3], [4] and [5] respectively.



A crude model for EAS Production of ionization e⁻ Time evolution of ionization e⁻ flux

・ロト ・ 同ト ・ ヨト ・ ヨ

Solution to Boltzmann equation obtained by MC simulation of all processes at work



Migration of e⁻ below the lowest excitation threshold of interest in about 1ns

The free-free approach Suppression effects

lonization electrons undergo quasi-elastic collisions with neutral molecules in the atmosphere

$$e^{\pm} + M \rightarrow e^{\pm} + M + \gamma$$

Photon production rate

$$r_{\gamma}(r,a,t,\nu) = \frac{n_m(a)}{\int_0^{T_e^{\max}} dT_e} \frac{\phi_{e,i}(r,a,T_e,t)}{\sigma_{\rm ff}(T_e,h\nu)} \sigma_{\rm ff}(T_e,h\nu)$$

where

- Density of molecules at altitude a
- Flux of ionization electrons
- Free-free cross section : for low-energy electrons and GHz photons $\sigma_{\rm ff}(T_e, h\nu) \rightarrow \sigma_{\rm ff}(T_e) = 1.211 \ 10^{-8} T_e \sigma_m(T_e)$ (Electron momentum transfer cross-section tables from [6])

The free-free approach Suppression effects

The emitted spectral power per volume unit at each point (r, a) :

$$\begin{aligned} \frac{d^2 P}{d\nu dV}(r, a, t) &= \frac{d}{d\nu} \left(h\nu \ r_{\gamma}(r, a, t)\right) \\ &= \frac{hc\rho_m^2(a)\mathcal{N}_A}{2AI_0} \left\langle \frac{dE}{dX} \right\rangle \tilde{\sigma} \ n_{e,p}(r, a) \exp(-t/\tau_{\text{att}}(a)) \end{aligned}$$

with

$$\tilde{\sigma} = \int_0^{T_e^{\text{max}}} dT_e \frac{I_0}{I_0 + T_e} \tilde{f}_0(T_e) \beta(T_e) \sigma_{\text{ff}}(T_e)$$

The effective cross section $\implies \tilde{\sigma} \simeq 5 \ 10^{-30} \ m^2$ (Nitrogen target)

ground level

The free-free approach Suppression effects

Semi analytical expression of the observable spectral intensity at any ground position $\mathbf{x}_{\mathbf{g}}$:

$$\Phi_{g}(\mathbf{x}_{g},t) = \int_{0}^{\infty} r dr \int_{0}^{2\pi} d\varphi \int_{0}^{\infty} da \frac{1}{4\pi R^{2}(r,\varphi,a)} \frac{d^{2}P}{d\nu dV}(r,a,t_{d}(t,r,\varphi,a))$$
Considerations:

Photons are emitted isotropically
Absorption effects found to be negligible

observation point x

< ロ > < 同 > < 回 > < 回 > < 回 > <

Do photons get absorbed within the interaction volume ? Absorption coefficient

- Defined as the relative attenuation per unit length of EM waves
- Derivation by making use of the detailed balance principle, absorption and spontaneous emission
- \Rightarrow Close to shower core and maximum of shower development $\alpha_{\nu}\simeq 10^{-19}~km^{-1}$
- \Rightarrow At GHz frequencies, the absorption is found to be negligible.

・ロト ・ 一日 ・ ・ 日 ・ ・ 日 ・ ・

 $\begin{array}{c} {\rm Introduction}\\ {\rm Ionization\ electrons\ along\ the\ shower\ track}\\ \mu-{\rm wave\ emission\ from\ MBR}\\ {\rm Simulation\ results}\\ {\rm Conclusion\ \&\ Outlook} \end{array}$

- Monte Carlo sampling of the observable spectral intensity in r and ϕ
- Spectral intensity as a function of time expected at different distances from the shower core at ground level, for a vertical shower with energy $10^{17.5}$ eV.



 \Rightarrow Close to the shower core, values are in the order or below from the ones measured of other sources of microwave radiations, such as geomagnetic effects.

Spectral intensity as a function of time expected at 10 km from the shower core at ground level, for a vertical shower with energy $10^{17.5}$ eV.



 \Rightarrow Values obtained at 10 km from the shower axis are a factor 25 less than expected when scaling beam measurements to air showers in reference [1].

 \Rightarrow Good sensitivity microwave detectors should detect the expected MBR intensity

Conlusion

- Detailed analytical calculation of the MBR spectral intensity has been undertaken
- Numerical solving of Boltzmann equation considering ionization, attachment and excitation of secondary electrons describe their time evolution in the plasma
- Calculated spectral intensity at ground gives a factor 25 lower than expected in [1]
- Still achievable detection with good sensitivity sensors
- Plasma dispersion effects are still to include in the calculation

< ロ > < 同 > < 回 > < 回 > < 回 > <

References

- 🔋 P. W. Gorham *et al.*, Phys. Rev. D78 (2008) 032007
- C. B. Opal *et al.*, J. Chem. Phys. 55 (1971) 4100
- 🔋 S. Nijdam *et al.*, J. Phys. D : Appl. Phys. 44 (2011) 455201
- D. Rapp, P. Englander-Golden, J. Chem. Phys. 43 (1965) 1464
- N. Kroll, K. N. Watson, Phys. Rev. A 5 (1972) 1883
- A. V. Phelps *et al.*, JILA cross-sections, http://jila.colorado.edu/~avp/collision_data/

イロト イポト イヨト イヨト