

Neutrino reconstruction for PINGU

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for the IceCube collaboration & PINGU

MANTS Meeting 14-15 October 2013 Garching, Germany





IceCube/DeepCore/PINGU

		At	tmospher	ic-v f	lux		
	Neutrino osc.[1,2]		Astrophysical-v sources				
	MSW effect (NMH)		Point sources - good angular resolution			Diffuse flux - good energy resolution	
							log (E/GeV)
(1) 1 GeV PINGU	2 DeepCore	1	3 TeV	4 IceCube	5	6 1 PeV



- * Cherenkov threshold $E_{th} > [n/(n^2-1)^{\frac{1}{2}}] \times m$ for $n_{ref} \sim 1.3$ (water/ice): $E_e > \sim 0.8$ MeV, $E_{\mu} > \sim 160$ MeV
 - [1] Antares collaboration, Phys.Lett. B714 (2012) 224 [1206.0645]
 [2] IceCube Collaboration, Phys. Rev. Lett. 111(2013), 081801 [1305.3909]



PINGU configurations



Closest distance between 2 strings: 20 m. (same for all strings in this configuration)

IceCube string #36 is indicated by a blue square. Red squares: 8 DeepCore strings Radius of the Circle, around IceCube String #36 is R=75 m. $\rho V_{PINGU} \sim 5 M ton$



Track reconstruction in IceCube

- ν_µ-CC and atmospheric-µ
 reconstruction in IceCube:
 "infinite track" approximation.
- PDF based reconstruction[1] includes best knowledge of calibrated detector and medium(ice) optical properties.
- The PDF is calculated with respect to "seed track" wich is obtained with first guess pattern recognition[2].



- [1] AMANDA Collaboration (J. Ahrens et al.), NIM. A524 (2004) 169 Muon track reconstruction and data selection techniques in AMANDA
- [2] IceCube Collaboration (M.G. Aartsen et al), arXiv:1308.5501 Improvement in fast particle track reconstruction with robust statistics



- Level 2 (L2) reconstruction: Standard IceCube/DeepCore reconstruction of the triggered/filtered events:
 - Hit cleaning algorithms
 - Various IceCube algorithms for track, vertex and energy reconstruction: SPE, MPE, FiniteReco, . . ., Monopod
- L2 reconstruction for DeepCore events where modified and updated for the PINGU MC data.
- Level2 reconstruction for PINGU includes additional algorithms:
 - SANTA [DESY]
 - IgelFit [Bonn U.]
 - HybridReco [PSU]



SANTA



 Used in current analysis of PINGU sensitivity [talk by A.Gross, MANTS-2013, 14/10/2013]



SANTA

 Energy and angular resolution obtained for PINGU MC events reconstructed with SANTA (direction) amd Monopod(energy)



A.Gross, presentations at ICRC-2013 (Rio de Janeiro) and VLVnT-2013 (Stokholm).



IgelFit

- Divides sphere around a vertex in N-segments/tracks (N=128 by default).
- In each segment IceCube PDF-based reconstruction is applied.
- Requires vertex coordinates as a "seed"

Currently studies: M. Day (UW) - for DeepCore data







- muon neutrino-CC event reconstruction with 8-parameters:
 - neutrino interaction vertex and time (x_0, y_0, z_0, t_0)
 - $\mu(v)$ direction: $\cos\Theta$, ϕ
 - μ -energy (defined from length I_{μ}) and cascade energy (E_x)
- Requires PDF / depends on the knowledge of ice properties.
 - Current PINGU MC simulations use SPICEMie.
 - PDFs are calculated from SPICE 1.
- Relatively slow, different minimization and LH-scan algorithms tested:
 - MINUIT algorithms
 - Markov Chain(MC): Vanilla Metropolis-Hastings MCalgorithm
 - MultiNest algorithm[1]
- [11 F. Feroz et al, arXiv< 0809,3437



MultiNest algorithm

João Pedro Athayde Marcondes de André @IceCube meeting, Munchen 2013



- Zenith angle resolution of less than 10° for most events
- Inner Gaussian contains about 57% of the events

Energy resolution of neutrino events reconstructed with HybridReco/Multinest (bottom left) and Monopod (bottom right)



Atmospheric v-flux and event selection

Atmospheric neutrino flux ("neutrino beam"):

M. Honda et al. Phys. Rev. D 83, 123001 (2011) [arXiv:1102.2688]

"Improvement of low energy atmospheric neutrino flux calculation using the JAM nuclear interaction model" HKKM fluxes: http://www.icrr.u-tokyo.ac.jp/~mhonda/nflx2011/index.html 6 sites: Frejus, Gran Sasso. INO, Kamioka, Soudan, Sudbury



- ν_µ-CC events ("track like") at low energy (E < ~20 GeV) will be contaminated with v-NC and electron and tau neutrino CC ("cascade") like events. PID-methods could help to separate these events.
- Method based on the "superluminal" hits for the separation of track and cascade type neutrino events studied in PSU. (talk by Ty. DeYoung)
- Another method is based on the variables from collider experiments for PID in the DeepCore/PINGU.
 Currently under development in NBI by M.Jørgensen.



Current PINGU MC data

- All PINGU MC data sets produced with GENIE.-GEANT-CLSIM* (/data/sim/PINGU/2012/triggered/GENIE-in-ice) are copied and stored at DESY. (* Ken Clark, GPU/WestGrid system)
- Largest statistics of simulated data: v15 (40 str., 60 PDOM) neutrino energy range: 1-80 GeV, E⁻²

	V _e	$ u_{\mu}$	v_{τ}
Simulated(x10 ⁸)	3.75	8.0	3.75
Triggered (x10 ⁶)	0.428	1.250	0.102



 All triggered events from PINGU v15 are reconstructed with "standard" reconstruction ("PINGU L2" including SANTA, Monopod)

Simulation volume for PINGU events



Summary and outlook

- Level2 reconstruction scripts for PINGU MC events based on standard IceCube/DeepCore algorithms and few new methods were set and tested. The PINGU reconstruction is working at different sites in USA/Canada and Europe.
- Current algorithms used in the reconstruction of v-direction (SANTA) and E_v (Monopod) provide sufficient angular and energy resolution for the NHM determination .
- New promising methods for the neutrino direction and energy reconstruction with improved resolutions were presented at the IceCube meeting last week.
- Work is in progress to adapt IceCube tables and algorithms developed for high energy neutrinos evens to lower energies..



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International Funding Agencies

Fonds de la Recherche Scientifique (FRS-FNRS) Fonds Wetenschappelijk Onderzoek-Vlaanderen (FWO-Vlaanderen) Federal Ministry of Education & Research (BMBF) German Research Foundation (DFG)

Deutsches Elektronen-Synchrotron (DESY) Inoue Foundation for Science, Japan Knut and Alice Wallenberg Foundation Swedish Polar Research Secretariat The Swedish Research Council (VR) University of Wisconsin Alumni Research Foundation (WARF) US National Science Foundation (NSF)